

**Tu02E-1**

# **A High LO-to-RF Isolation E-band Mixer with 30 GHz Instantaneous IF Bandwidth in 90nm CMOS**

**Wei-Chieh Ma<sup>1</sup>, Chau-Ching Chiong<sup>2</sup>,  
Yun-Shan Wang<sup>1</sup>, Huei Wang<sup>1</sup>**

**<sup>1</sup>Graduate Institute of Communication Engineering,  
National Taiwan University, Taipei, Taiwan**

**<sup>2</sup>Academia Sinica Institute of Astronomy and Astrophysics  
(ASIAA), Taipei, Taiwan**

# Outline

- Motivation
- Paper Survey
- Schematic
- Circuit Design
  - Bias and Transistor Size
  - Butterworth LPF
  - LC Resonators
- Simulation and Measurement
- Comparison to Reported Mixers
- Conclusion
- References

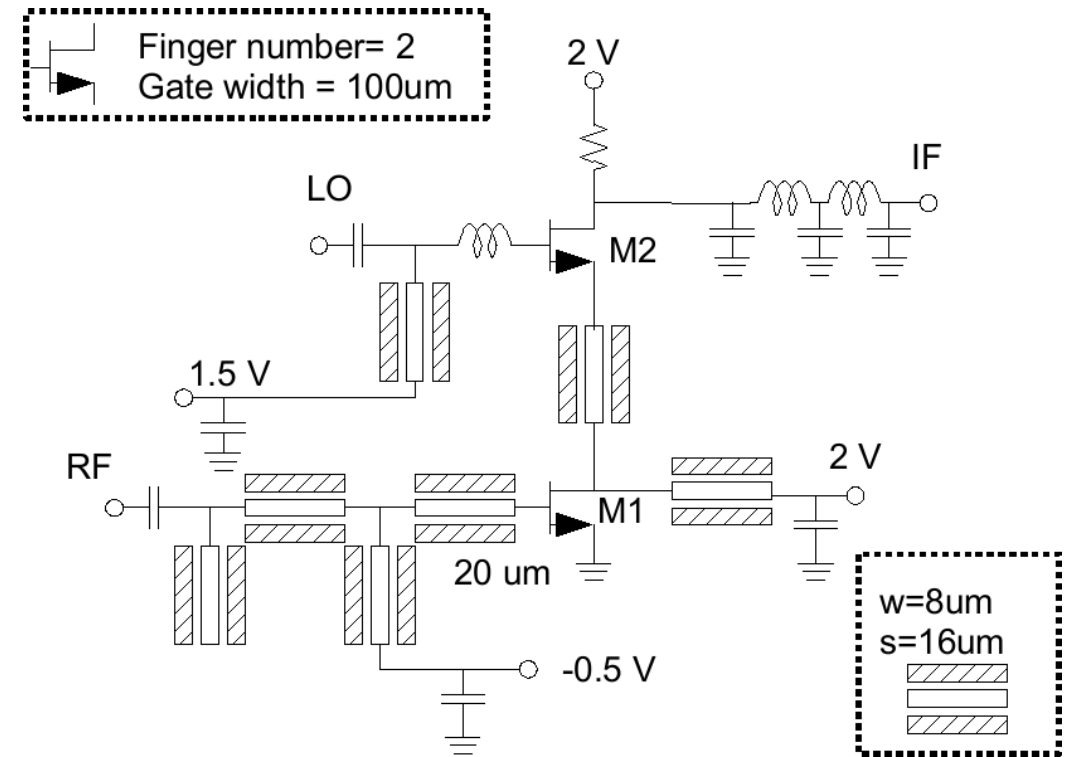
- **ALMA2030 Development Roadmap**
  - Goals: Explore origins of planets, galaxies, and chemical complexity
- **Atacama Large Millimeter/submillimeter Array (ALMA)**
  - High-sensitivity
  - High-resolution

- ALMA receivers specification
  - Low-noise performance
    - Increase sensitivity of observation
  - Instantaneous bandwidth
    - Improve efficiency of observation

# Paper Survey

- A 24-48 GHz Cascode HEMT Mixer with DC to 15 GHz IF Bandwidth for Astronomy Radio Telescope

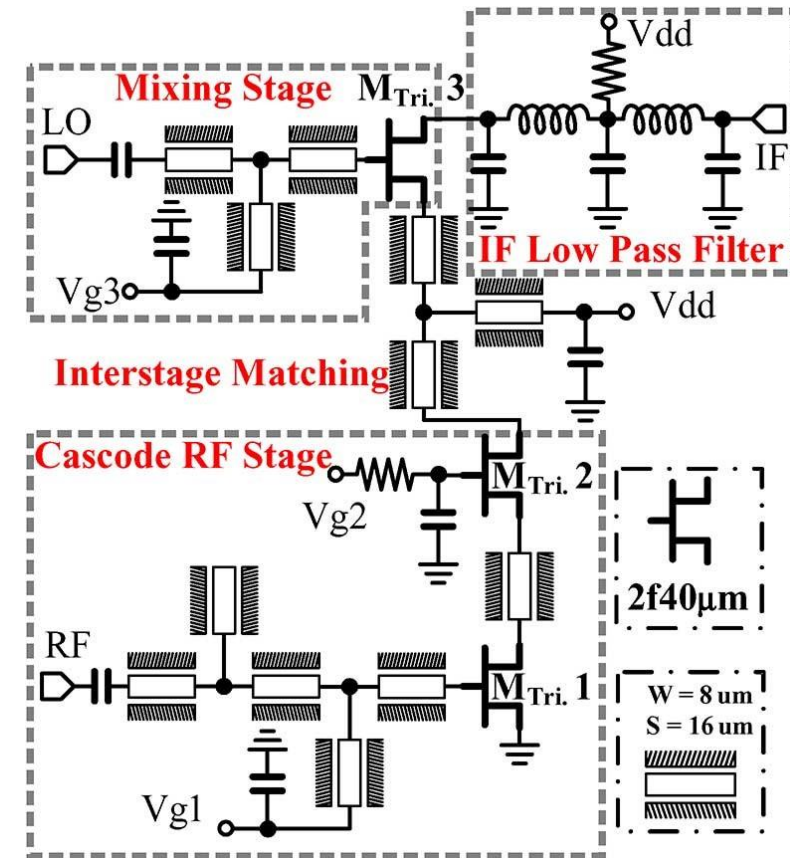
- Cold-biased mixing stage
- Common-source RF stage
- Poor LO-to-RF isolation
- IF frequency: DC-15 GHz



[1] Z. -M. Tsai, J. -C. Kao, K. -Y. Lin and H. Wang, "A 24–48 GHz cascode HEMT mixer with DC to 15 GHz IF bandwidth for astronomy radio telescope," 2009 European Microwave Integrated Circuits Conference (EuMIC), 2009, pp. 5-8.

- A W-band High LO-to-RF Isolation Triple Cascode Mixer With Wide IF Bandwidth

- Cold-biased mixing stage
- Cascode RF stage
- Degraded IP1dB
- IF Frequency: DC-24 GHz

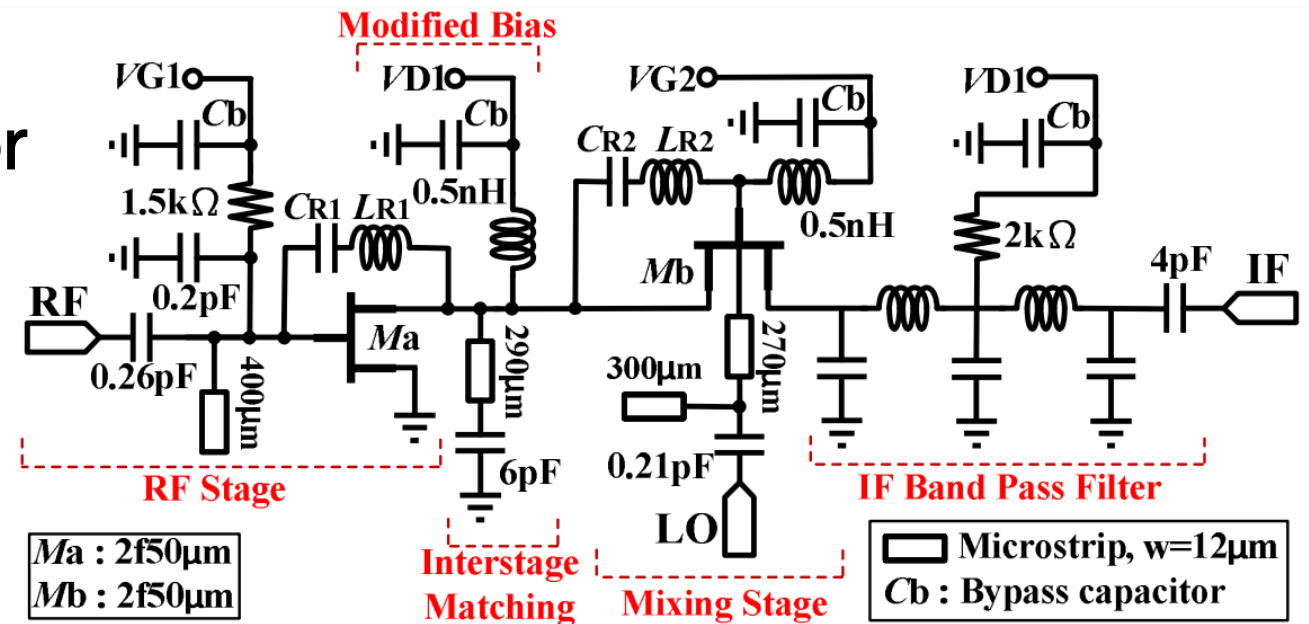


[2] J. -C. Kao, K. -Y. Lin, C. -C. Chiong, C. -Y. Peng and H. Wang, "A W-band High LO-to-RF Isolation Triple Cascode Mixer With Wide IF Bandwidth," in IEEE Transactions on Microwave Theory and Techniques, vol. 62, no. 7, pp. 1506-1514.

# Paper Survey

## • A High LO-to-RF Isolation 34-53 GHz Cascode Mixer for ALMA Observatory Applications

- Cold-biased mixing transistor
- Common-source RF stage
- **LC resonators**
- IF frequency: 3-13 GHz



[3] C. -N. Chen, Y. -H. Lin, Y. -C. Chen, C. -C. Chiong and H. Wang, "A High LO-to-RF Isolation 34–53 GHz Cascode Mixer for ALMA Observatory Applications," 2018 IEEE MTT-S International Microwave Symposium (IMS), 2018, pp. 686-689.

# Paper Survey

- [1] Poor LO-to-RF isolation
- [2] Sacrifice IP1dB for LO-to-RF isolation
- [3] Chosen structure

Ref.	Mixing Stage	RF Stage	Embedding	IF Bandwidth	LO-to-RF Isolation	IP1dB
[1]	Cold-biased	Common-source	N/A	😊	😞	😊
[2]	Cold-biased	Cascode	N/A	😊	😊	😞
[3]	Cold-biased	Common-source	LC resonators	😊	😊	😊

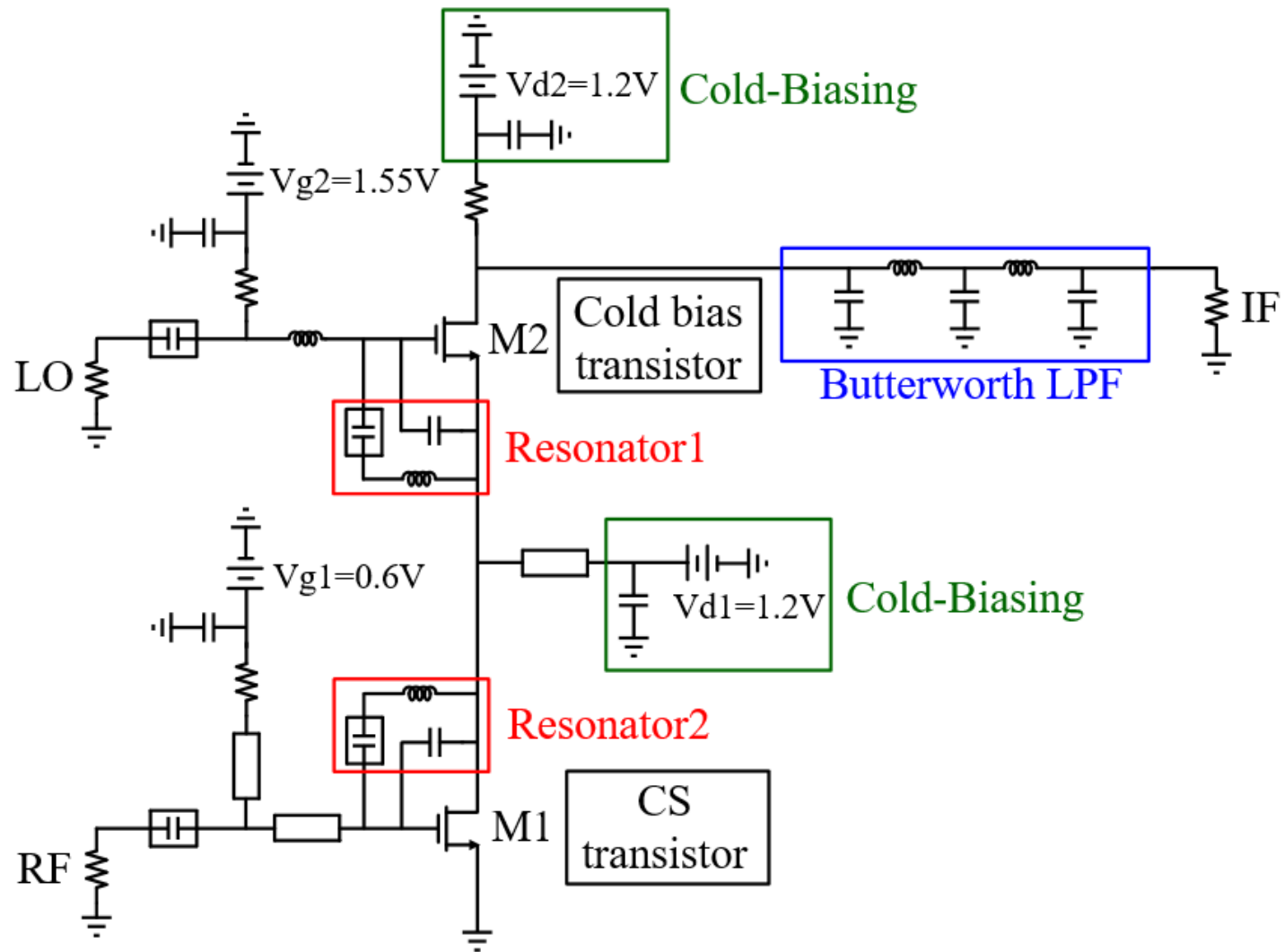


# Paper Survey

- 90nm CMOS process (Compared to GaAs pHEMT)
  - Lower parasitic capacitance
    - More suitable for wide IF bandwidth design
  - Multi-layer flexibility
  - Compact layout
  - Lower cost

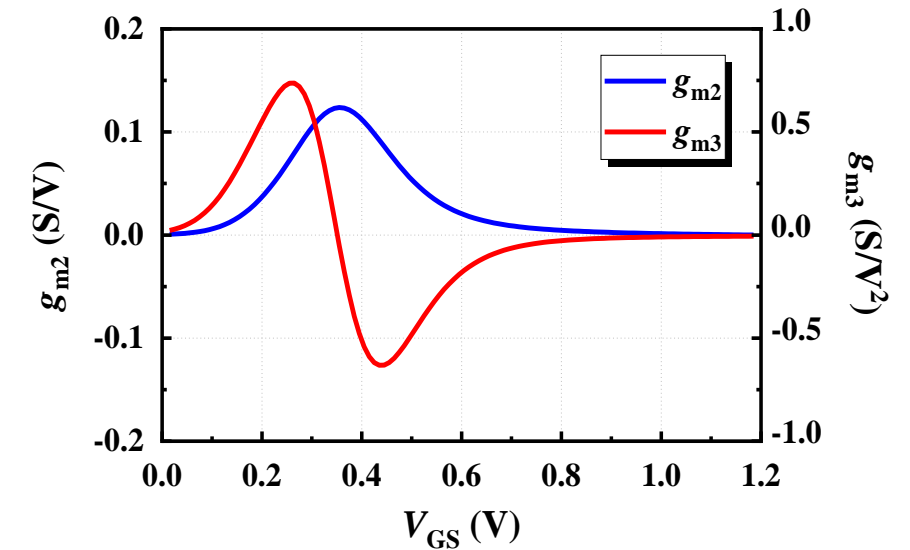
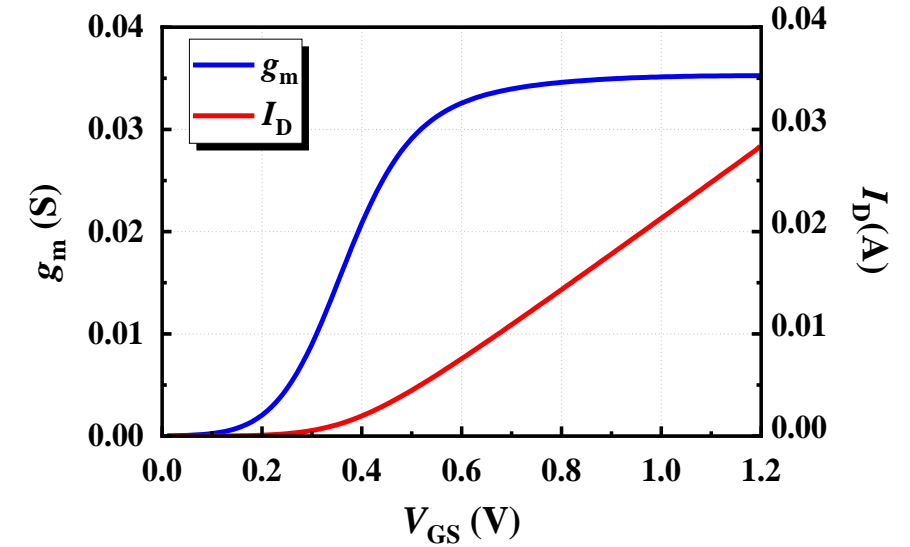
# Schematic

- **Cold-Biasing Technique**
  - Extend IF bandwidth
- **Butterworth LPF**
  - Improve LO-to-IF isolation
  - Improve RF-to-IF isolation
- **LC Resonators**
  - Improve LO-to-RF isolation



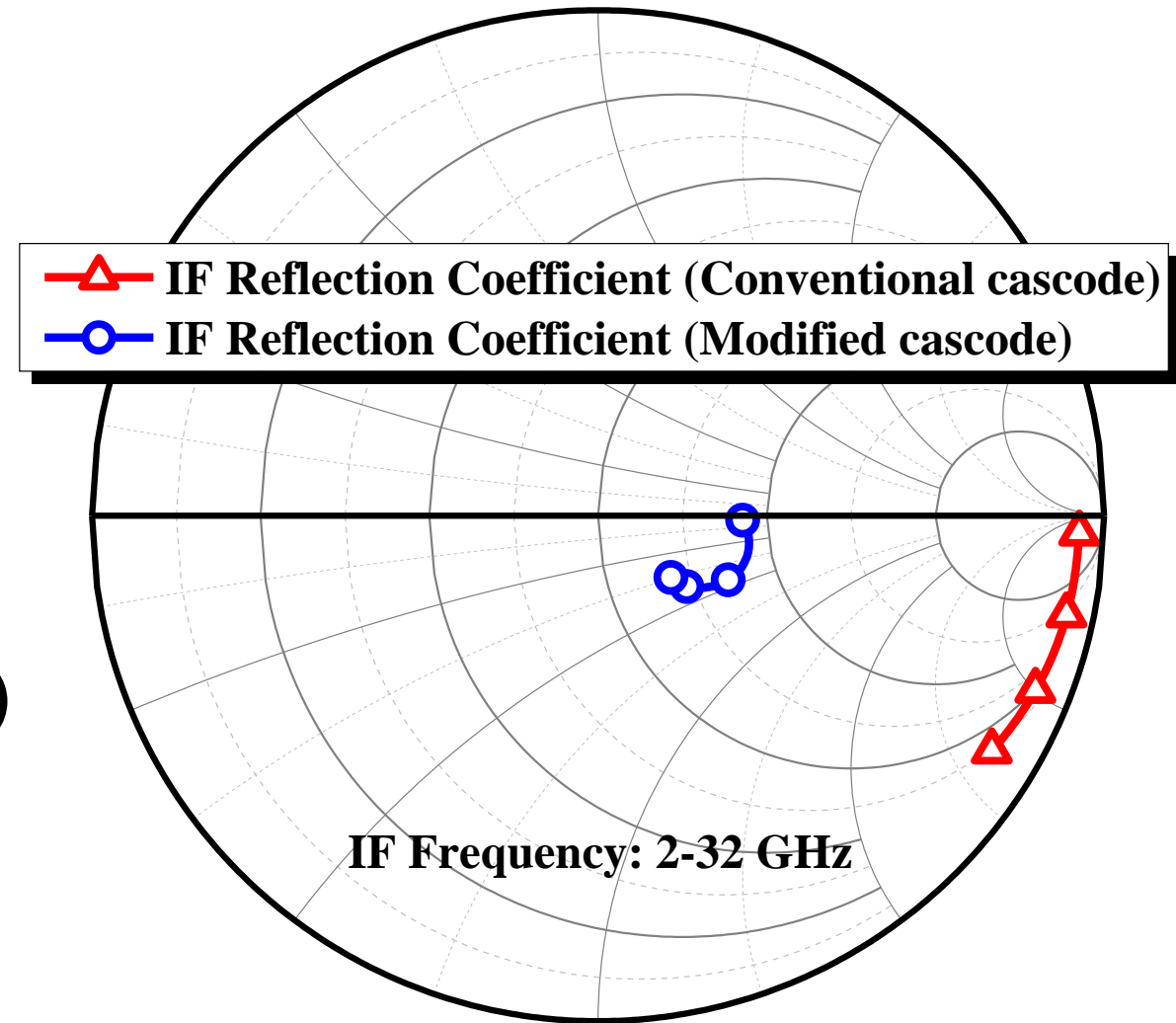
# Circuit Design (Bias)

- RF Stage ( $V_{GS}=0.6$  V)
  - Trade-Off between  $g_m$  and  $P_{DC}$ 
    - $V_{GS} \nearrow : g_m \nearrow, P_{DC} \nearrow$
- Mixing Stage ( $V_{GS}=0.35$  V)
  - Maximize second-order transconductance
    - Good conversion gain
  - Minimize third-order transconductance
    - Good linearity



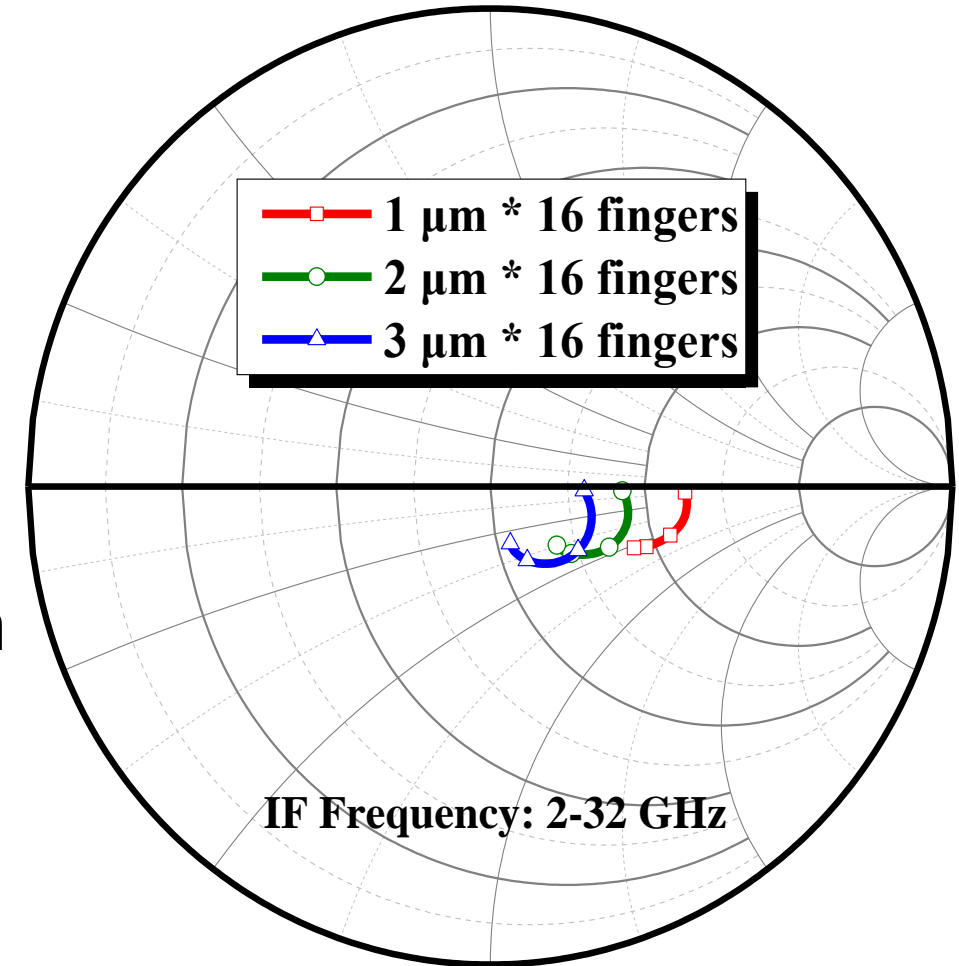
# Circuit Design (Bias)

- **Conventional Cascode**
  - Common gate mixing stage
  - Output impedance
    - Higher
    - More frequency-dependent
- **Modified Cascode**
  - Cold-biased mixing stage ( $V_{DS}=0$ )
  - Output impedance
    - Lower (closer to 50 Ohm)
    - Less frequency-dependent



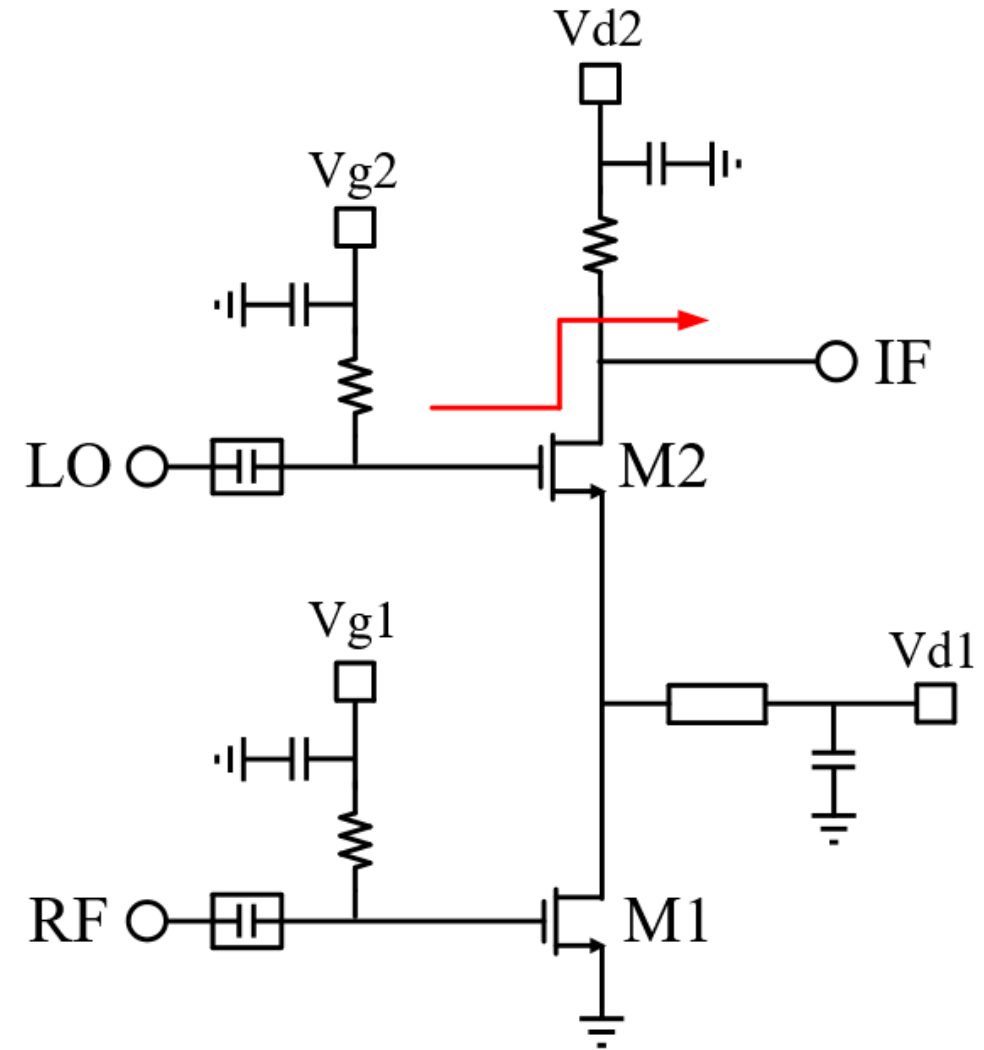
# Circuit Design (Transistor Size)

- RF Stage ( $2\ \mu\text{m} * 16$  fingers)
  - Trade-off between  $g_m$  and  $P_{DC}$ 
    - Size  $\nearrow$  :  $g_m \nearrow$ ,  $P_{DC} \nearrow$
- Mixing Stage ( $2\ \mu\text{m} * 16$  fingers)
  - Trade-off between CG and bandwidth
    - Size  $\nearrow$  : CG  $\searrow$ , bandwidth  $\nearrow$
  - Output impedance close to 50 Ohm



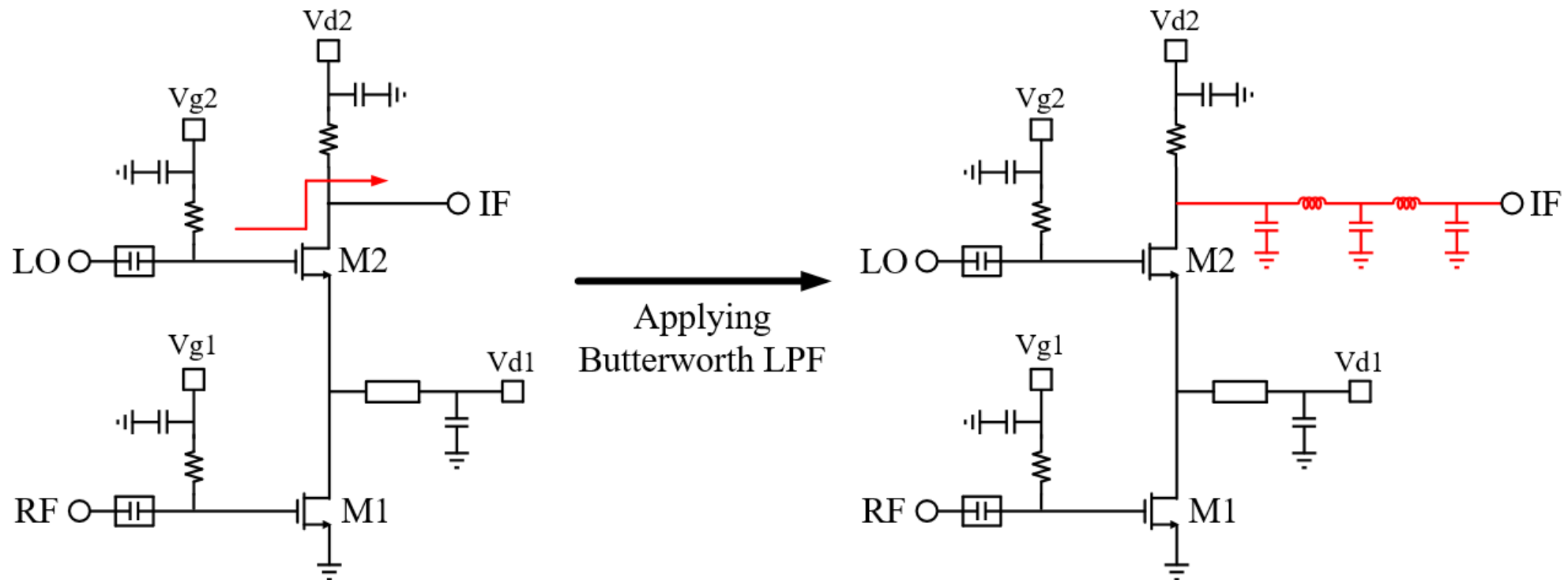
# Circuit Design (Butterworth LPF)

- LO-to-IF Leakage
  - Saturate IF amplifier



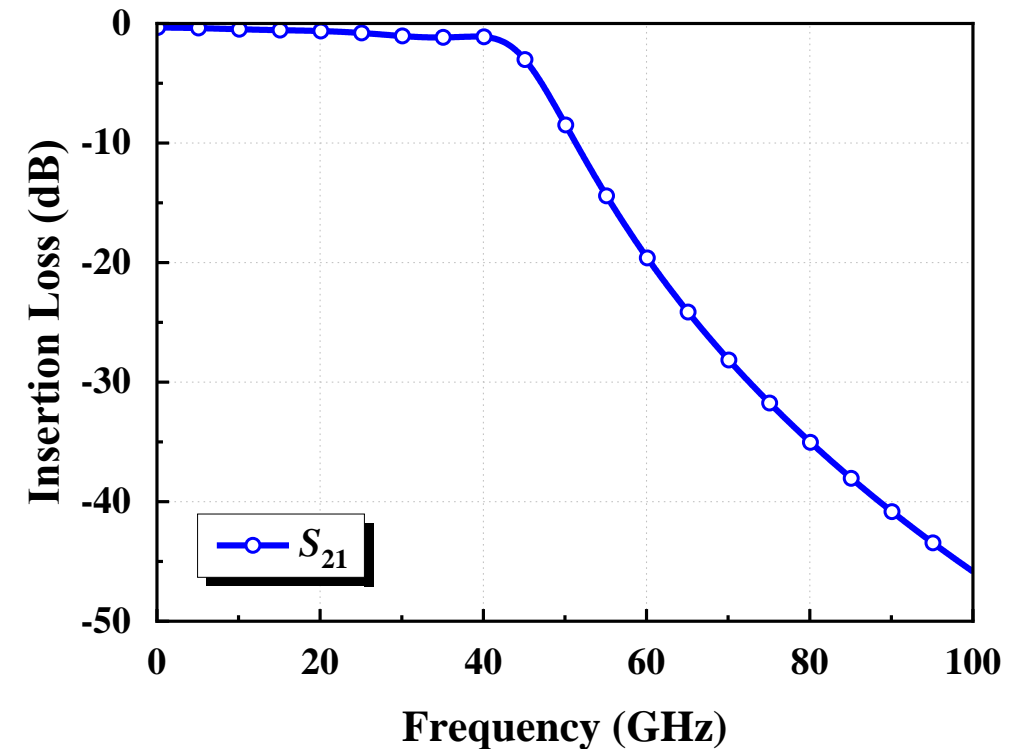
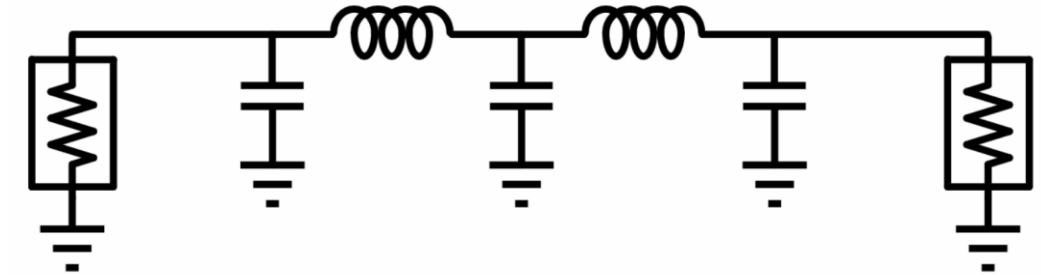
# Circuit Design (Butterworth LPF)

- Butterworth LPF
  - Pass IF signal (slightly matching for gain flatness)
  - Block LO and RF signal



# Circuit Design (Butterworth LPF)

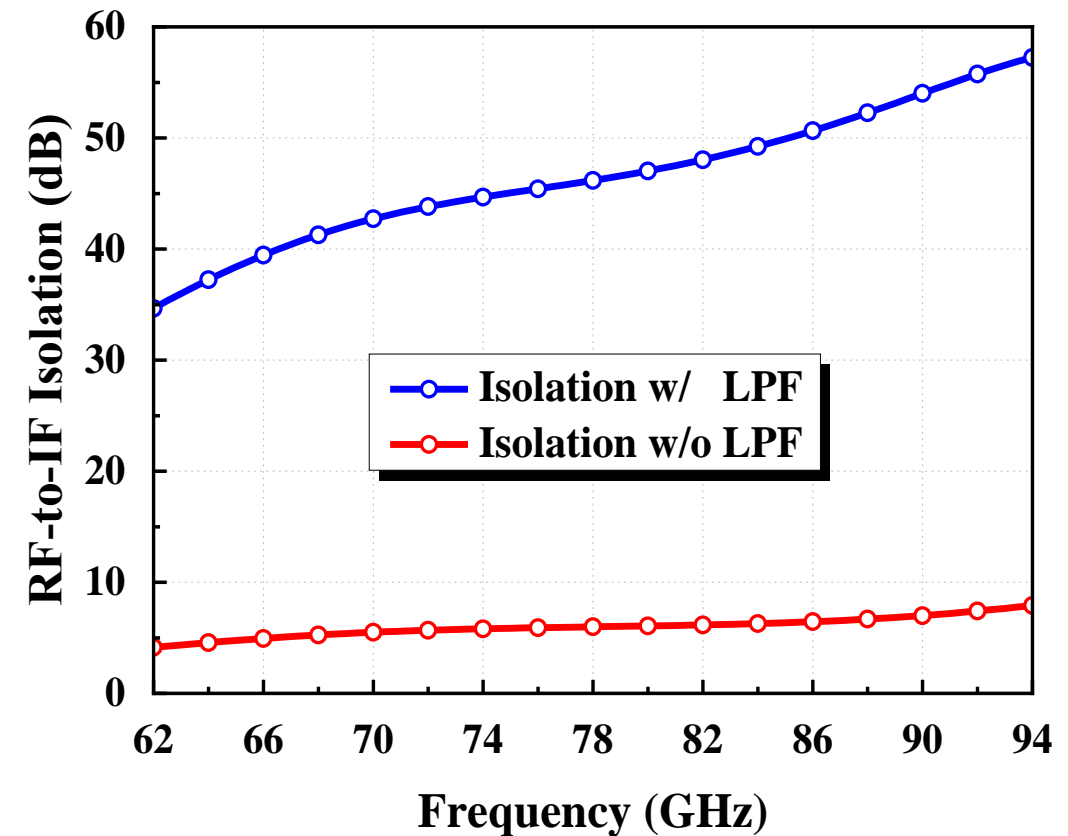
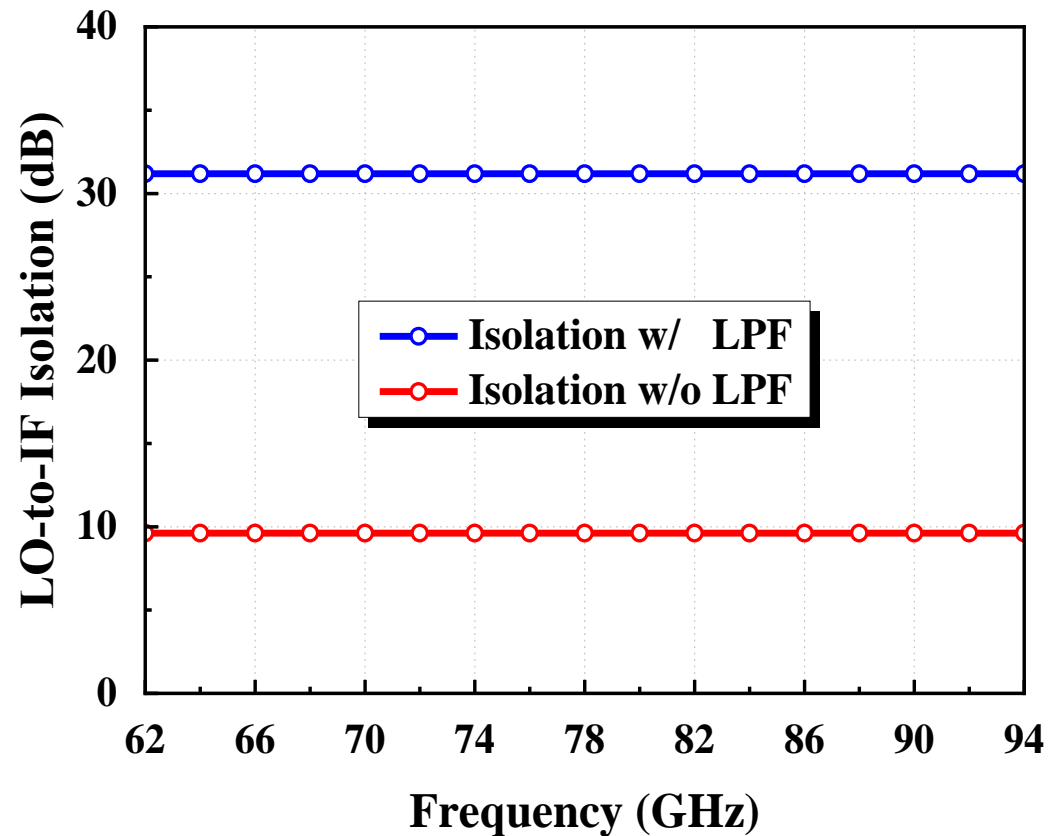
- 5<sup>th</sup> order shunt-first LPF
  - Passband (DC-32 GHz)
    - Insertion loss lower than 1 dB
  - Stopband (60-92 GHz)
    - Insertion loss higher than 20 dB

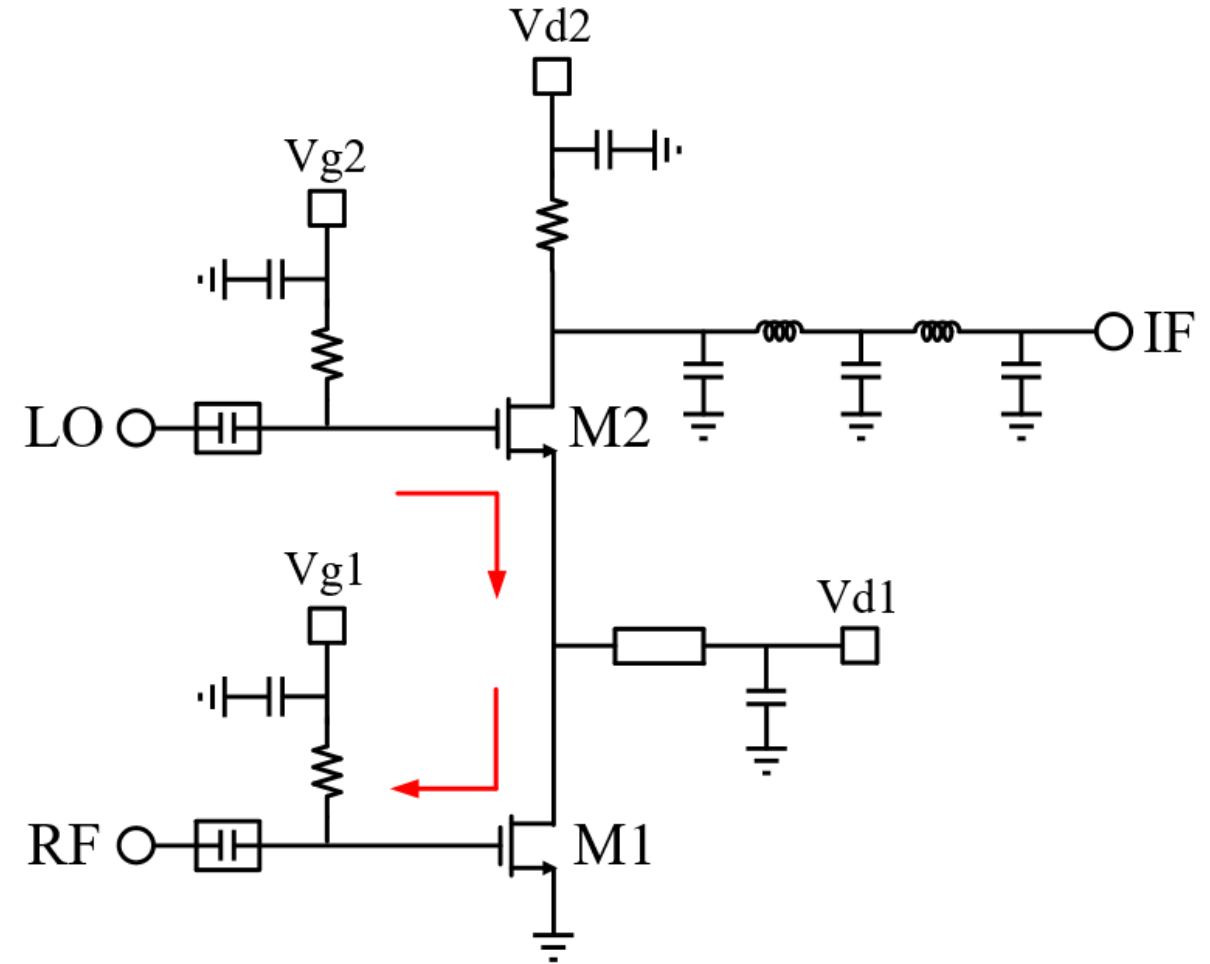




# Circuit Design (Butterworth LPF)

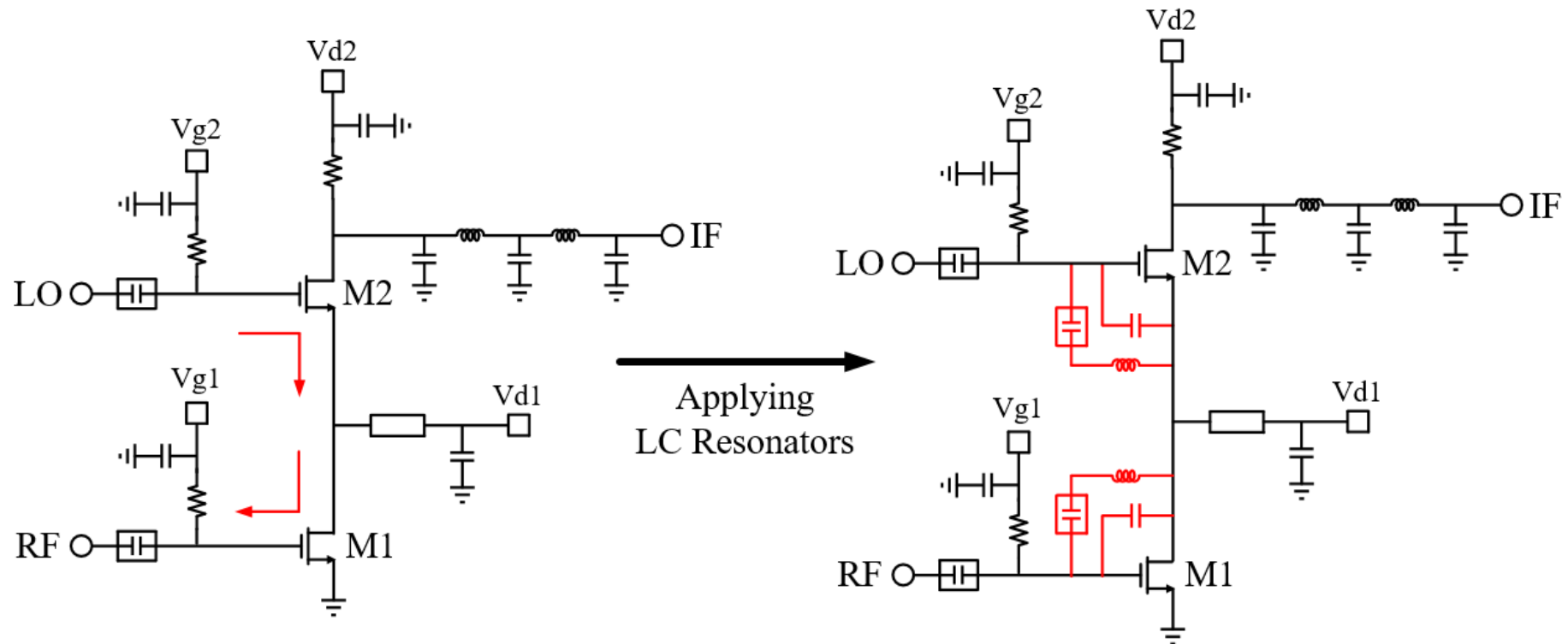
- Improvement in LO-to-IF isolation and RF-to-IF isolation





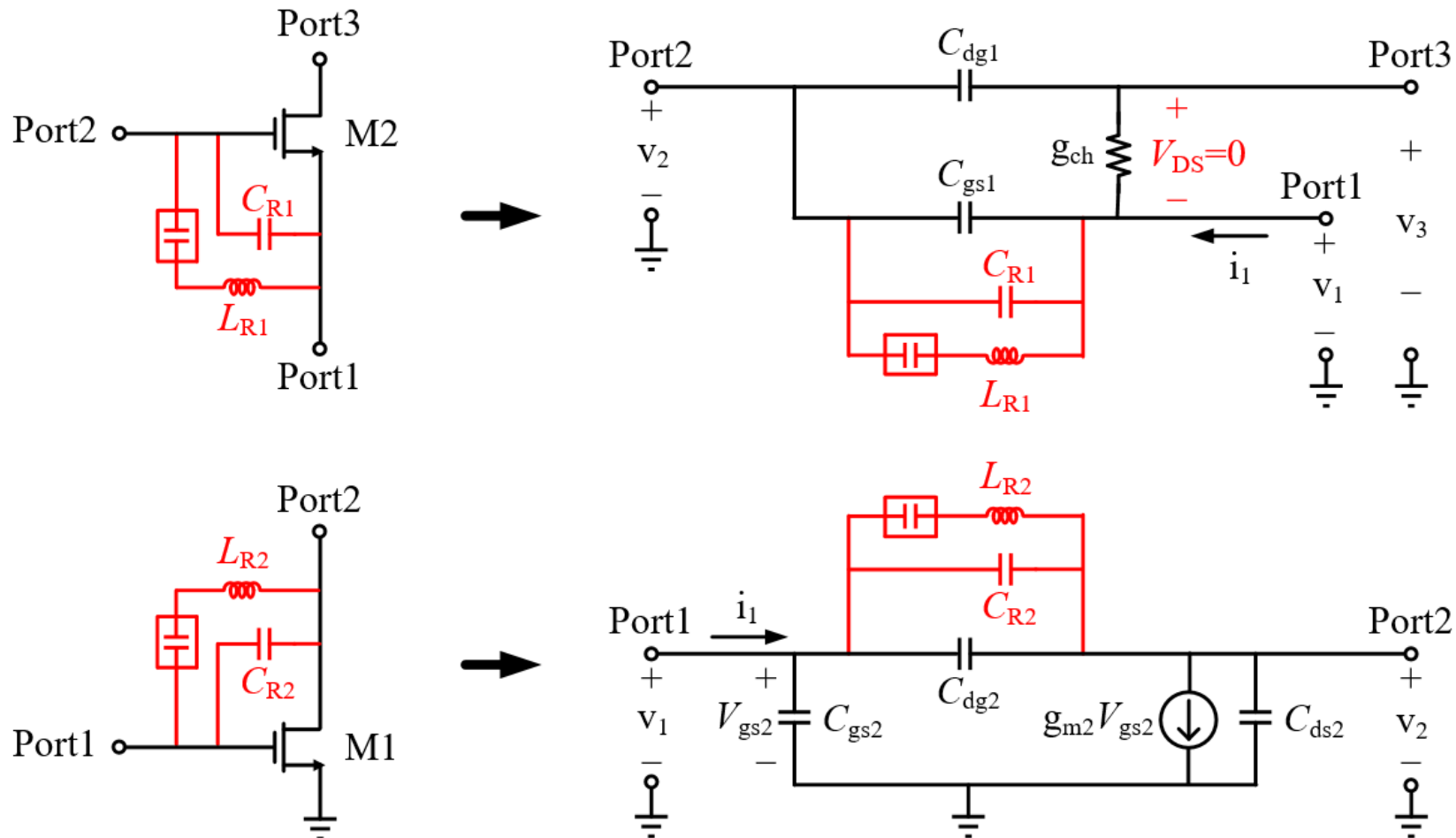
# Circuit Design (LC resonators)

- LC resonators
  - Parallelled LC pairs resonating at LO frequency (60 GHz)
  - Parasitic capacitor included ( $C_{gs}$  of mixing stage and  $C_{dg}$  of RF stage)



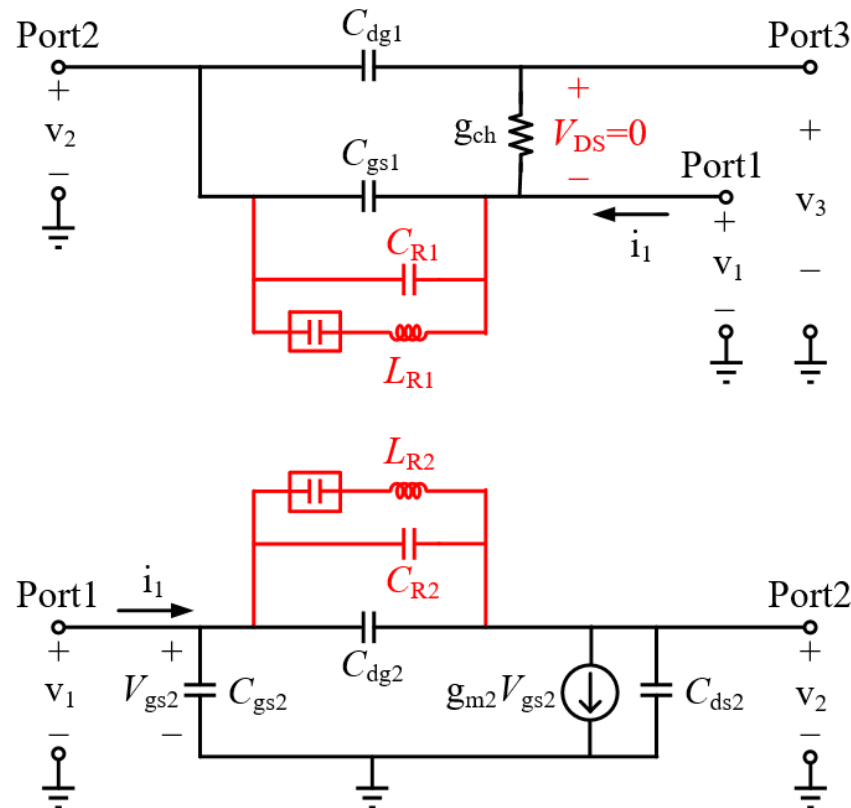
# Circuit Design (LC resonators)

- Small signal equivalent circuit



# Circuit Design (LC resonators)

- Admittances between ports of LO-to-RF leakage path
  - Both need to be minimized!

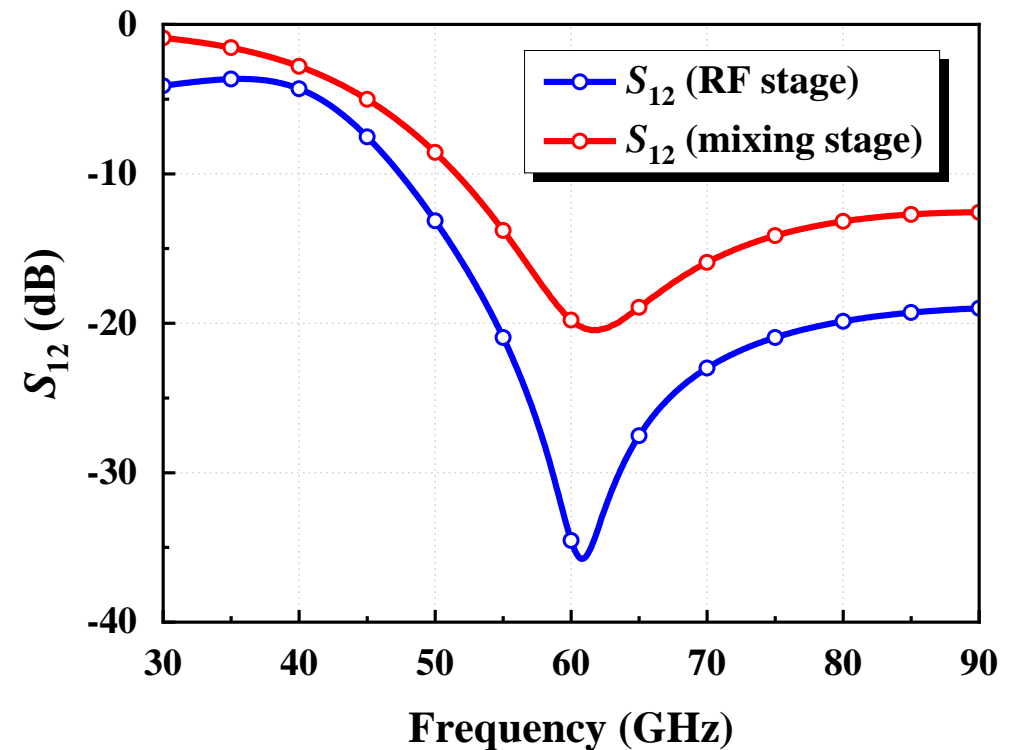
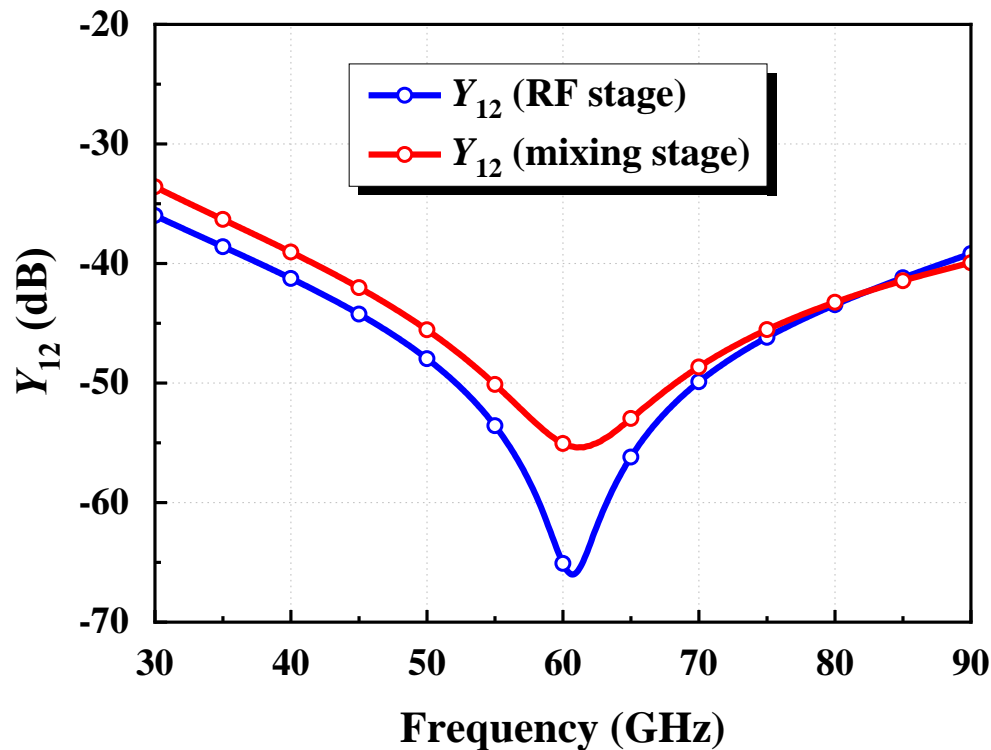


$$Y_{12} = \frac{i_1}{v_2} \Big|_{v_1=0}^{v_3=0} = -j2\pi f \left( C_{R1} + C_{gs1} - \frac{1}{4\pi^2 f^2 L_{R1}} \right)$$

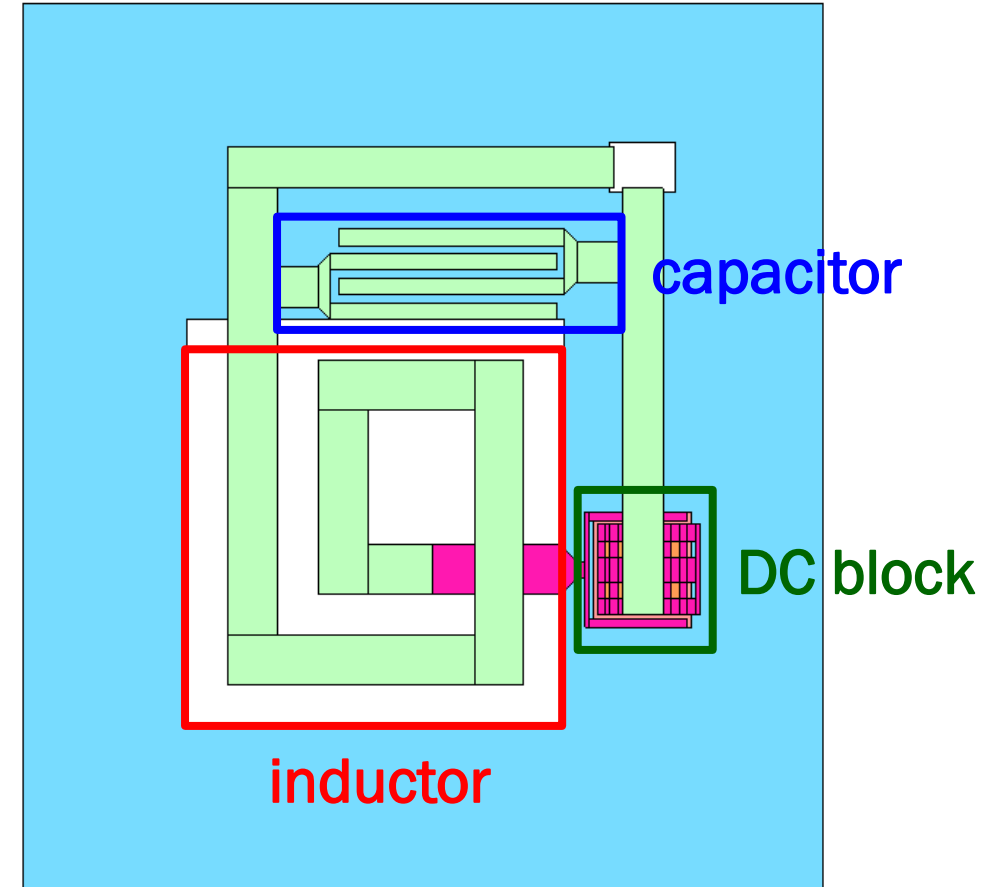
$$Y_{12} = \frac{i_1}{v_2} \Big|_{v_1=0} = -j2\pi f \left( C_{R2} + C_{dg2} - \frac{1}{4\pi^2 f^2 L_{R2}} \right)$$

# Circuit Design (LC resonators)

- Inductance and Capacitance
  - Minimize  $Y_{12}$  at 60 GHz
    - LC Resonance frequency: 60 GHz

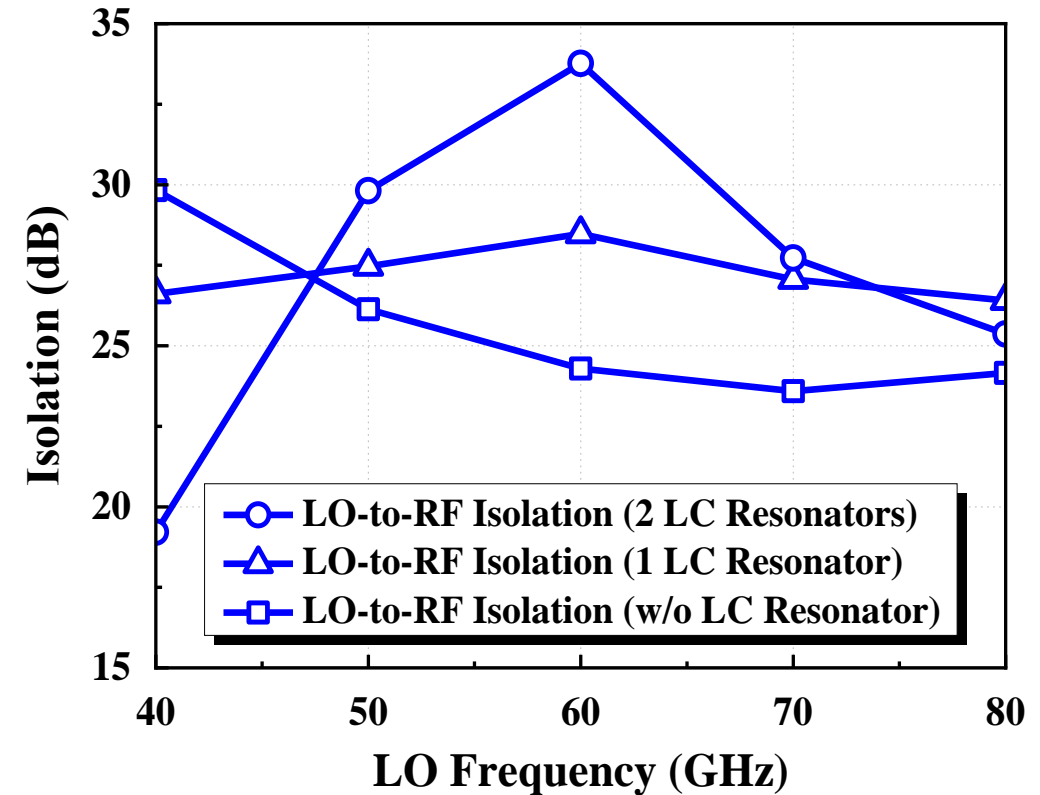


- Layout consideration
  - Spiral inductor
  - Edge-coupled MOM capacitor
    - Reduce impact of process variation
  - DC block
    - Large capacitor with negligible reactance



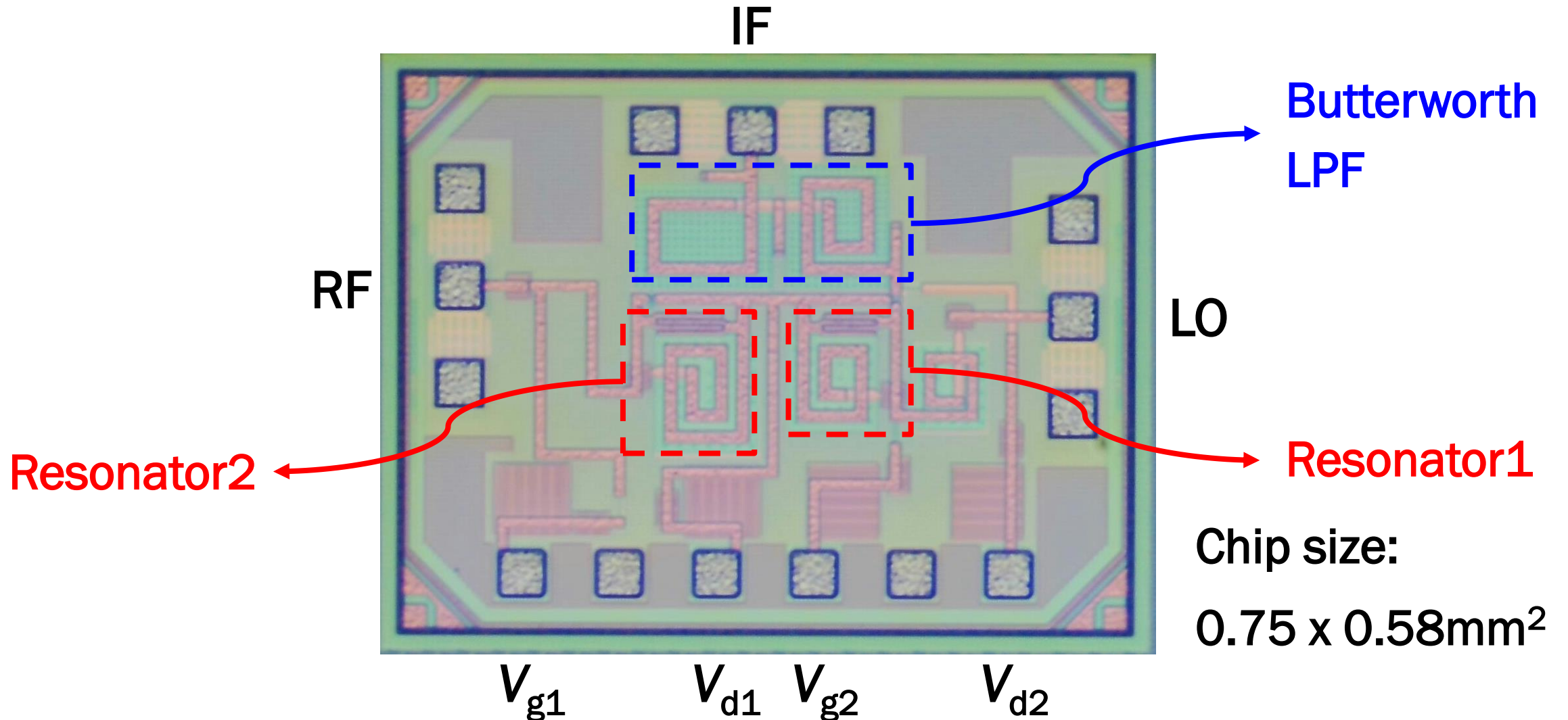
# Circuit Design (LC resonators)

- Without LC Resonator
  - 24 dB LO-to-RF Isolation
    - Signal interference and DC offset
- With 2 LC Resonators
  - 34 dB LO-to-RF Isolation
    - Meeting system requirement



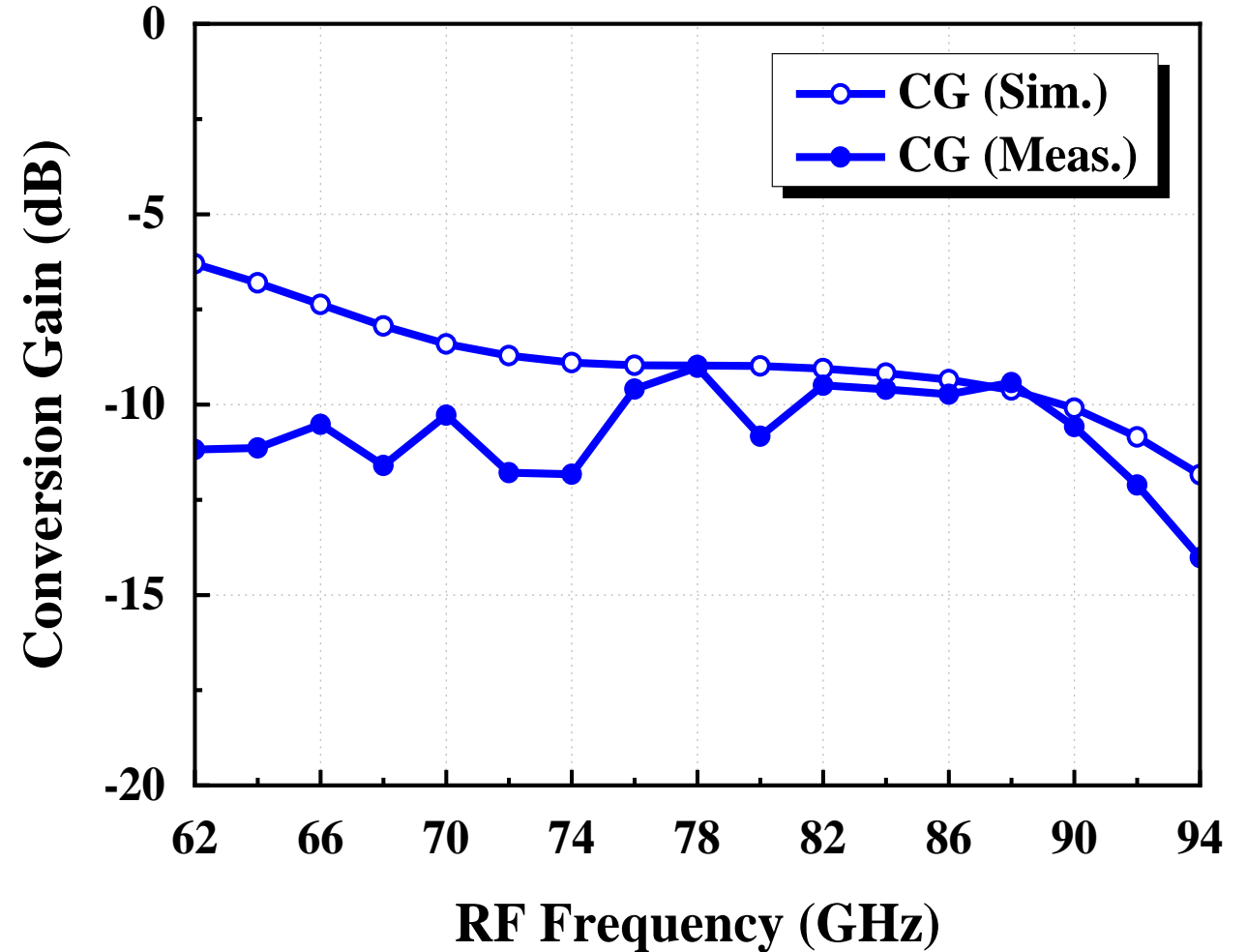


# Simulation and Measurement



- Conversion Gain
  - -9 to -12 dB
- 3 dB IF Frequency
  - 2 to 32 GHz

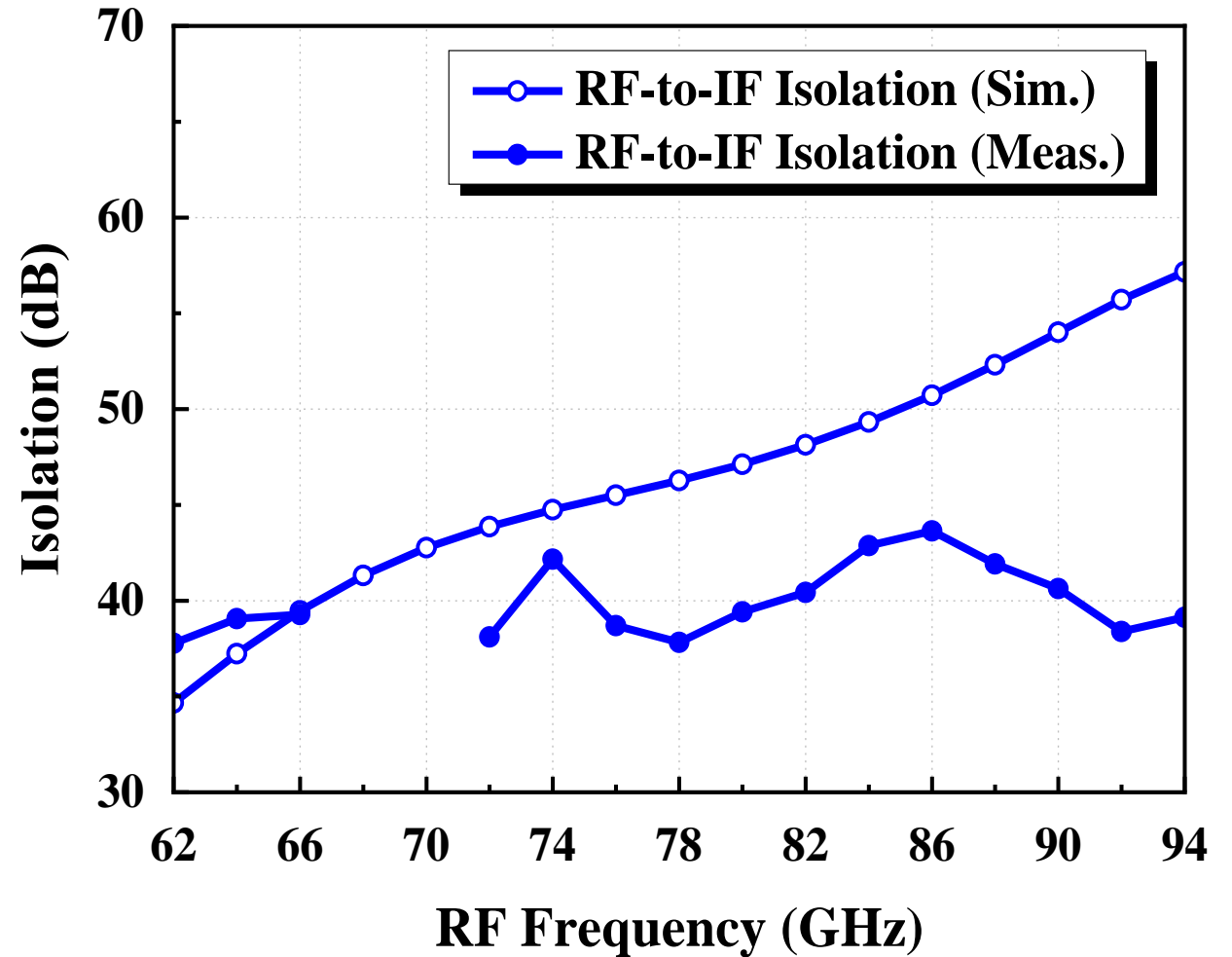
(With 8-dBm 60-GHz LO signal)



- LO-to-RF Isolation
  - 40 dB
- LO-to-IF Isolation
  - 31 dB

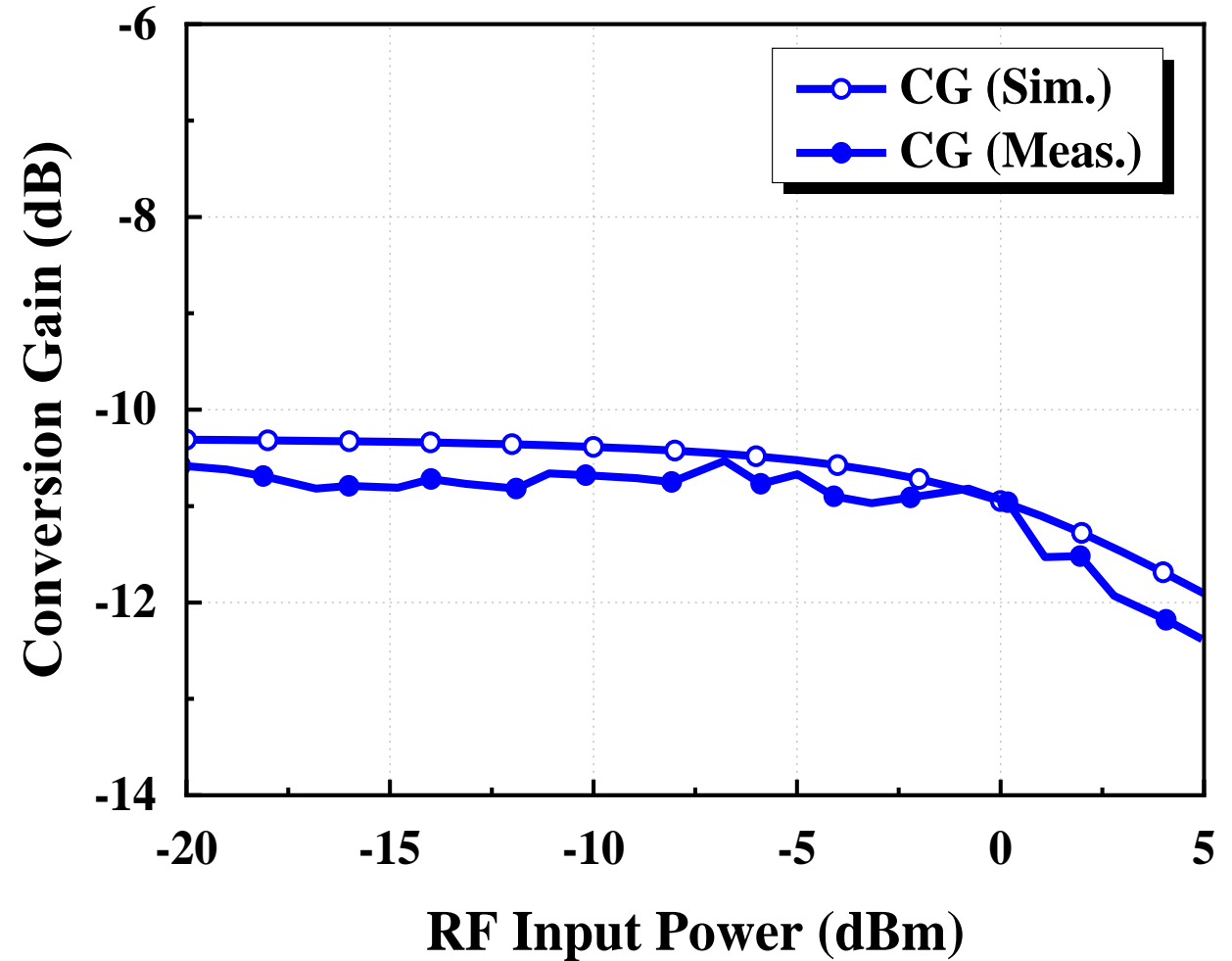
(Single LO frequency: 60 GHz)
- RF-to-IF Isolation
  - > 38 dB

(RF frequency: 62-94 GHz)



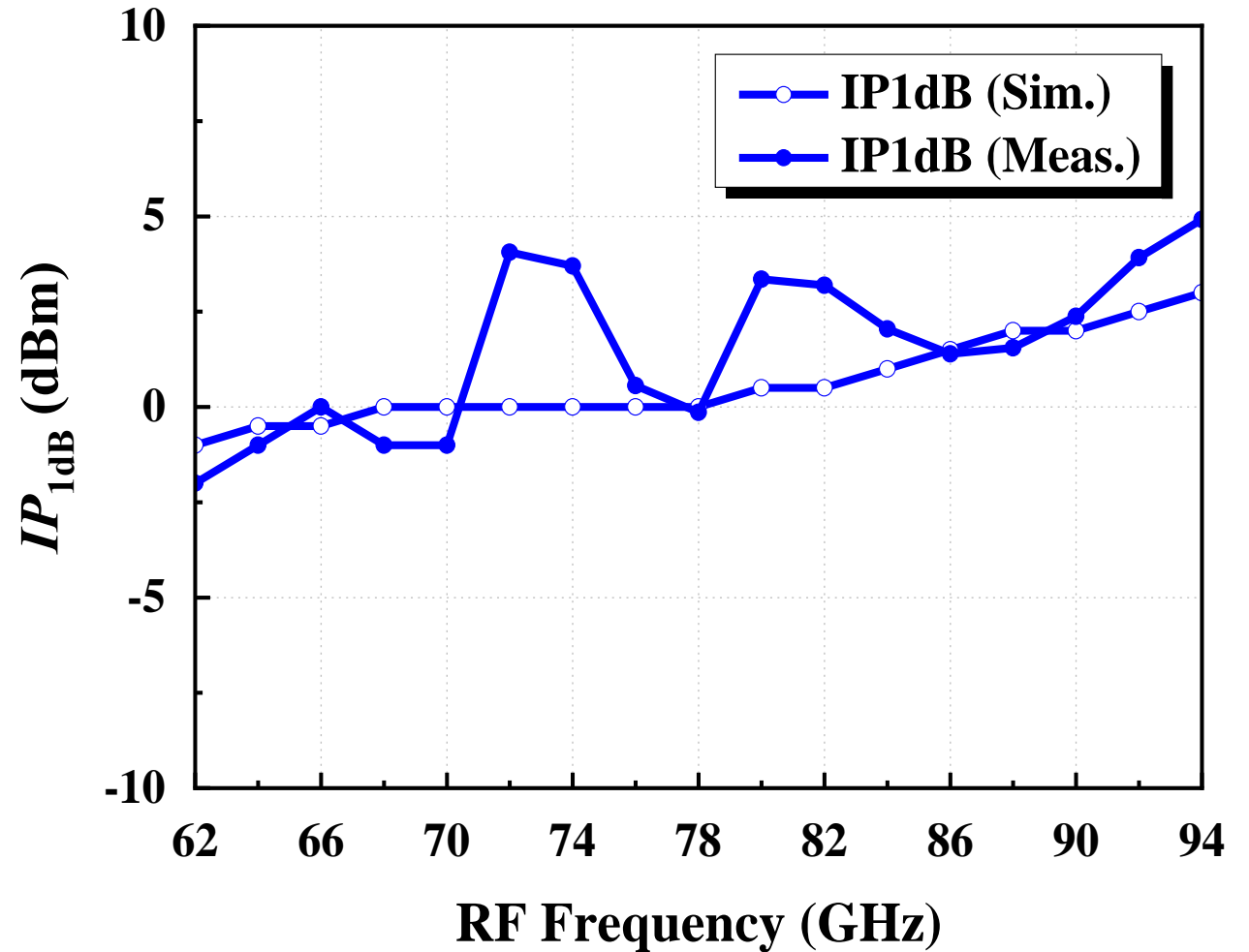
- IP1dB at 90 GHz  
– 2 dBm

(With 8-dBm 60-GHz LO signal)



- IP1dB in Full E-band  
– -2 to 4 dBm

(With 8-dBm 60-GHz LO signal)



# Comparison to reported mixers

Ref.	Process	Topology	RF freq (GHz)	IF freq (GHz)	CG* (dB)	$IP_{1dB}$ (dBm)	LO-to-RF ISO# (dB)	$P_{dc}$ (mW)	LO power (dBm)	Area (mm <sup>2</sup> )
[1]	0.15- $\mu$ m GaAs pHEMT	Modified cascode	27-47	DC-15	0	N/A	N/A	N/A	4	1.5
[2]	0.15- $\mu$ m GaAs pHEMT	Modified triple cascode	75-120	DC-24	-10 - -17	N/A	41.5	24	7	1
[3]	0.1- $\mu$ m GaAs pHEMT	Modified cascode	34-53	3-13	-2--4	-2--4	37	36	0	2.3
[4]	0.25- $\mu$ m GaAs pHEMT	Star mixer	17-34	0.1-13	-6.6--15.2	N/A	22.6	0	14	0.8
[5]	0.15- $\mu$ m GaAs pHEMT	Ring mixer	40-50	DC-10	-9.2--11.9	N/A	20	0	15	1.2
[6]	0.15- $\mu$ m GaAs pHEMT	Subharmonically pumped diode	75-105	DC-21	-4.7	N/A	15	0	11	2
[7]	GaAs	Double balanced	18-50	DC-21	-8.7	9	39	0	12-22	16
[8]	90nm CMOS	Fundamental drain/gate pumped	30-90	DC-16	-7.2--13.7	2	30.2	0.6	4.2	0.4
<b>This work</b>	<b>90nm CMOS</b>	<b>Modified cascode</b>	<b>62-92</b>	<b>2-32</b>	<b>-9-12</b>	<b>-2-4</b>	<b>40.8</b>	<b>8.1</b>	<b>8</b>	<b>0.4</b>

\*Conversion gain, # LO-to-RF isolation

# Conclusion

- Cold-Biasing Technique
  - Extend Instantaneous IF Bandwidth
    - 2-32 GHz 3-dB IF frequency
- LC Resonators
  - Improve LO-to-RF Isolation
    - 40 dB LO-to-RF Isolation
- Acceleration in Astronomical Observation
  - The observation is two to three times faster than before



- [1] Z. -M. Tsai, J. -C. Kao, K. -Y. Lin and H. Wang, "A 24–48 GHz cascode HEMT mixer with DC to 15 GHz IF bandwidth for astronomy radio telescope," 2009 European Microwave Integrated Circuits Conference (EuMIC), 2009, pp. 5-8.
- [2] J. -C. Kao, K. -Y. Lin, C. -C. Chiong, C. -Y. Peng and H. Wang, "A W-band High LO-to-RF Isolation Triple Cascode Mixer With Wide IF Bandwidth," in IEEE Transactions on Microwave Theory and Techniques, vol. 62, no. 7, pp. 1506-1514.
- [3] C. -N. Chen, Y. -H. Lin, Y. -C. Chen, C. -C. Chiong and H. Wang, "A High LO-to-RF Isolation 34–53 GHz Cascode Mixer for ALMA Observatory Applications," 2018 IEEE MTT-S International Microwave Symposium (IMS), 2018, pp. 686-689.
- [4] Y.-A. Lai, C.-N. Chen, S.-H. Huang, and Y.-H. Huang, "Compact Double-balanced Star Mixers with Novel Dual 180° Hybrids," in Proc. IEEE 11th Int. Conf. Solid-State Integr. Circuit Technol., Oct. 2012, pp.1-4.
- [5] Z. Chen, X. Jiang, W. Hong, and J. Chen, "A Q-band doubly balanced mixer in 0.15um GaAs PHEMT technology," in IEEE Int. Wireless Symp., Mar. 2014, pp. 1-4.
- [6] Y.-J. Hwang, H. Wang, and T.-H. Chu, "A W-band subharmonically pumped monolithic GaAs-based HEMT gate mixer," IEEE Microw. Wireless Compon. Lett., vol. 14, no. 7, pp. 313–315.
- [7] Markimicrowave, "GaAs MMIC Double Balanced Mixer," MM1-1850HSM datasheet, May. 2019 [Revised July. 2019].
- [8] Y.-C. Wu, C. -C. Chiong and H. Wang, "A novel 30–90 GHz singly balanced mixer with broadband LO/IF," 2016 IEEE MTT-S International Microwave Symposium (IMS), 2016, pp. 1-4.



# Thank you for listening!